

Gaussian Splatting Lucas-Kanade

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Introduction

Background

- •Dyna mic 3D scene reconstruction from monocular video is limited by sparse viewpoints.
- •Dynamic Gaussian splatting requires a ccurate geometry for high-quality rendering.

Challenges

- •Limited parallax leads to incorrect warp fields and unstable motions
- •Depth and optical flow priors are noisy and fail to regularize dynamics.

Related Works

•Learne d warp fields via MLPs or offsets. •Structural cues (depth, flow) applied, but insufficient for reliable motion modeling.

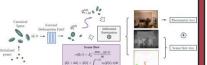
Contributions

- •De rive analytical velocity fields from the warp field via Luc as-Ka nade adaptation.
- •Enable continuous time integration without extra networks.
- •Improve motion accuracy and tractability over data-driven methods.

Failure Modes

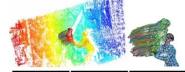


Regularizing Warp Field



The warp field can be regularized by suppressing the irregularities in the motion Jacobians for smoother predictions.

Underlying motion fields







Without regularization, the Gaussians could overcompensate for photometric quality with inaccurate motions.

Experiment Results

Experiment Setup

- •Evaluated on five dynamic scene datasets with varying camera motion (DyCheck, Dynamic Scene, Plenoptic Video, HyperNeRF, DAVIS 2017).
- •Achieves state-of-the-art results under static and dyna mic cameras, validated quantitatively and qualitatively.

Reconstructed glass geometries



The underlying geometry is more complete with the warp field regularization

Comparisons with baseline methods



Quantitative Comparisons

Experiments on Dynamic Scenes dataset

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method	PSNR ↑	Playground SSIM †	LPIPS ‡	PSNR ↑	Balloon! SSIM ?	LPIPS‡	PSNR ↑	Balloon2 SSIM †	LPIPS ‡	PSNR ↑	Umbrella SSIM↑	LPIPS
NSFF	24.69	0.889	0.065	24.36	0.891	0.061	30.59	0.953	0.030	24.40	0.847	0.088
NR-NeRF	14.16	0.337	0.363	15.98	0.444	0.277	20.49	0.731	0.348	20.30	0.526	0.315
Nerfles	22.18	0.802	0.133	23.36	0.852	0.102	24.91	0.864	0.089	24.29	0.803	0.169
(w flow)	22.39	0.812	0.109	24.36	0.865	0.107	25.82	0.899	0.081	24.25	0.813	0.123
Flow-sup, NeRF	16.70	0.597	0.168	19.53	0.654	0.175	20.13	0.719	0.113	18.00	0.597	0.148
Deformable GS	24.82	0.646	0.343	22.40	0.833	0.137	24.19	0.818	0.153	22.35	0.711	0.186
4DGaussians	21.39	0.776	0.204	24.48	0.849	0.144	24.72	0.801	0.219	21.29	0.560	0.332
Ours	26.34	0.756	0.384	26.38	0.848	0.133	25.89	0.911	0.151	23.82	0.746	0.176

DyCheck dataset and ablated results

Method		Apple	Spin	Block	Teddy	Paper Winds	Method	PSNR	SSIM	LPIP
NSFF Marties 4D Gaussians Ours Ours - pose refined		17.54 17.57 15.41 16.03 22.61	18.38 15.49 14.41 16.71 24.09	16.61 16.88 11.28 15.46 23.27	13.65 13.57 12.36 13.60 17.78	17.34 18.67 15.60 17.41 19.91	Deformable GS w/ flow Deformable GS w/ depth Deformable GS w/ both Deformable GS Ours	20.97 21.17 23.94 24.82 26.34	0.690 0.673 0.702 0.646 0.756	0.31 0.25 0.28 0.34 0.18

Warp Field Uncertainty



The covariance of warp Jacobians can be used to estimate motion uncertainties over time.

Conclusion

We derive analytical velocity fields to regularize dynamic Gaussian splatting, improving motion a ccuracy and reconstruction quality from limited views. Future work explores 3D tracking from casual captures and dynamic scene rendering in the wild.

Acknowledgement

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